

ASPECTS OF DE SITTER SPACETIMES



Dionysios Anninos

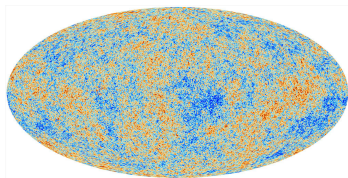


THE
ROYAL
SOCIETY

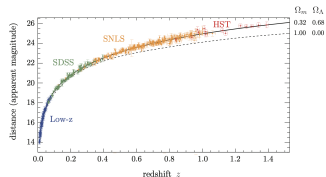
Eurostrings, Gijón-Xixón, April 24th 2023

SHAPE OF THE LARGE SCALE UNIVERSE

Cosmological & astrophysical observations indicate our Universe experiences epochs of accelerated expansion



Planck Collaboration



Left Cosmic microwave background (CMB) inhomogeneities (2003 – onwards)

Right Type Ia supernova measurements (1998 – onwards)

Quantum effects play a role in both quasi de Sitter epochs



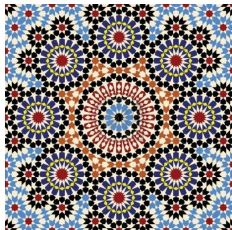
Inflationary hypothesis Quantum fluctuations of light fields, including gravitational field, are responsible for refined patterns in the CMB

Guth-Linde-Albrecht,Steinhart-Starobinsky-Mukhanov,Chibishov-Pi,Hawking-...

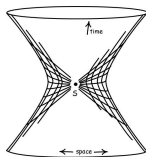
Dark energy Quantum contributions to Λ \rightsquigarrow cosmological constant problem

Nernst-...-Lenz-...-Zeldovich-...-Weinberg-...

GEOMETRY & SYMMETRIES



$$\frac{ds^2}{\ell^2} = -dT^2 + \cosh^2 T d\Omega_d$$



- No natural notion of spatial/null infinity or S -matrix.
- Isometries = $SO(d + 1, 1)$.
- At \mathcal{I}^+ $SO(d + 1, 1)$ acts as the conformal group on S^d

ETERNAL INFLATION & CONFORMAL FIELD THEORY

Eternal Inflation has symmetries of 3d Euclidean CFT

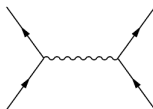
-Hull-...-Strominger-Witten-...-Maldacena-...

- Late time profiles of field operators transform as conformal operators
- At the free level, associate profiles to unitary irreps of $SO(d + 1, 1)$

Harish Chandra-...-Hirai-...-Higuchi-...- Basile,Bekaert,Boulanger-...-Letsios-...

- Correlations at \mathcal{I}^+ satisfy CFT Ward identities

...-Pimentel,Maldacena-...-Bzowski,McFadden,Skenderis-...



- **4pt functions** tree-level & h.s. exchange...

...-Mata,Raju,Trivedi-...-D.A.,Denef,Monten,Sun-...-ArkaniHamed,Baumann,Joyce,Pueyo,Lee,Pimentel-Sleight,Taronna-Bonifacio,Goodhew,Joyce,Pajer,Stefanyszyn-...

HARTLE-HAWKING & CONFORMAL FIELD THEORY

Late-time Hartle-Hawking wavefunction governed by 3d CFT

...-Maldacena-...-Hartle,Hawking,Hertog-...-D.A.,Anous,Freedman,Konstantinidis-...-Chakraborty,Chakravarty,Godel,Paul,Raju-

$$\implies \boxed{\Psi[g] = Z_{\text{CFT}}[g]}$$



BULK CALCULATION:

$$f(\mathbf{k}_i) = \int \frac{d\eta}{\eta^4} \prod_i G(\mathbf{k}_i, \eta)$$



BOUNDARY CALCULATION:

$$f(\mathbf{k}_i) = \langle \prod_i \mathcal{O}(\mathbf{k}_i) \rangle_{\text{CFT}}$$

Applications to cosmology ...-Leigh,Larsen,van der Schaar-...-McFadden,Skenderis-...

DS/CFT?

A complete dS/CFT conjecture should likely address:

- Gauge invariant observables at \mathcal{I}^+ & measure for $|\Psi[g]|^2$
- Relatedly – Hilbert space and microscopic operator content at \mathcal{I}^+
- Unitarity* & locality \leftrightarrow reality & locality properties of CFT
- Relatedly – Analogue/replacement of state-operator correspondence?

* Cosmological bootstrap developments:

...-Sengor,Skordis-Anous,Skulte-Benincasa-ArkaniHamed,Baumann,Joyce,Pueyo,Lee,Pimentel-Sleight,Taronna-Goodhew,Jazayeri,Pajer-Penedones,Hogervorst,SalehiVasiri-Gorbenko,Komatsu-...

- Recent progress in ‘finite time’ $\Psi[g]$ and T^2 deformations

AraujoRegado,Khan,Wall-...

- Issues can be concretely addressed in exotic higher-spin theories

...-D.A.,Hartman,Strominger-...-D.A.,Denef,Harlow-D.A.,Denef,Monten,Sun-...Neiman-...-Strominger,Cotler-...

COSMOLOGICAL EVENT HORIZON



DE SITTER'S STATIC UNIVERSE

Λ dominated world \rightsquigarrow observers surrounded by **cosmological event horizon**



This is the cosmological no hair conjecture

Hawking, Moss-Wald...

Physics classically **inaccessible** beyond the de Sitter horizon

DE SITTER HORIZON THERMODYNAMICS

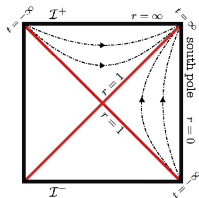
STATIC GEOMETRY: $ds^2 = - \left(1 - \frac{r^2}{\ell^2}\right) dt^2 + \left(1 - \frac{r^2}{\ell^2}\right)^{-1} dr^2 + r^2 d\Omega^2$

Temperature: $T = \frac{\hbar c}{2\pi\ell}$

Figari, Höegh Krohn, Nappi-Gibbons, Hawking-...

Entropy: $S_0 = \frac{A_{\text{horizon}}}{4G} = \frac{3\pi c^3}{\Lambda \hbar G}$

Bekenstein-Hawking...Gibbons, Hawking...Banks, Fischler-Parikh, Verlinde-Susskind...Dong, Silverstein, Torroba-...



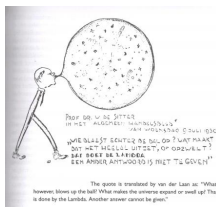
In our world $S_0 \sim 10^{122} \gg S_{\text{BH}} \sim 10^{100} \gg S_{\text{matter}} \sim 10^{90}$

EUCLIDEAN PATH-INTEGRAL

Gibbons & Hawking, more concretely, postulate the **quantum** entropy

$$\mathcal{S} \equiv \log \mathcal{Z}_{\text{grav}} = \log \int \mathcal{D}g e^{-S_E[g_{ij}, \Lambda]}$$

for fields on a **sphere** topology



\mathcal{S} is **diffeomorphism & field redefinition invariant** quantity

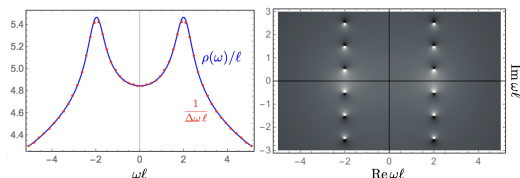
MATTER CONTRIBUTIONS

Thermal gas for arbitrary species content

$$\log \mathcal{Z}_{\text{matter}} = \sum_{\text{species}} \int_{\mathbb{R}^+} d\omega \rho(\omega) \log \frac{e^{-\frac{\beta\omega}{2}}}{1 - e^{-\beta\omega}}$$

where $\rho(\omega) = \rho_{\text{bulk}}(\omega) - \rho_{\text{edge}}(\omega)$ is group theoretically fixed

D.A., Deneff, Law, Sun-...



$\rho_{\text{edge}}(\omega)$ stem from edge modes (for $s \geq 1$) localised at the horizon

Buidovich, Polikarpov-...-Kabat-...-Donnelly, Wall-...

MACROSCOPIC PRECISION FORMULAE

3+1 dimensional gravity – one loop result on S^4

$$\mathcal{S} = \mathcal{S}_0 - 5 \log \mathcal{S}_0 - \left(\frac{331}{90} + \frac{8}{3} \right)_{\text{b+e}} \log \frac{\ell^2}{\ell_{\text{ref}}^2} + 15.3655 + \dots$$

...-Christensen,Duff-...-Volkov,Wipf-...-D.A.,Denef,Law,Sun-...-Benedetti,Casini-David,Mukherjee-...

2+1 dimensional gravity – all loop result on S^3

$$\mathcal{S}_{3d} = \mathcal{S}_0 - \log \frac{\mathcal{S}_0}{4\pi} + 2 \log \sinh \frac{4\pi^2}{\mathcal{S}_0}$$

Witten-...-Carlip-Guadagni,Tomassini-...-Castro,Lashkari,Maloney-...-Gukov,Mariño,Putrov-...-D.A.,Denef,Law,Sun-D.A.,Harris-...

LOWER DIMENSIONAL METHODS



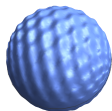
2D QUANTUM GRAVITY

General structure Gibbons-Hawking path-integral + 2d CFT

$$\log \mathcal{Z}_{\text{grav}} = \vartheta + \left(\frac{24}{(\sqrt{c-1} - \sqrt{c-25})^2} - 1 \right) \log \frac{1}{\Lambda} + f(c)$$

Knizhnik, Polyakov, Zamolodchikov-...-Distler, Kawai-David-...

For $c \rightarrow +\infty$ (timelike Liouville) one can tame geometric fluctuations



Euclidean dS_2 semiclassical saddle + systematic all loop conjecture for $f(c)$

Polchinski-...-Giribet-...-D.A., Bautista, Beatrix Mühlmann-Mühlmann-...

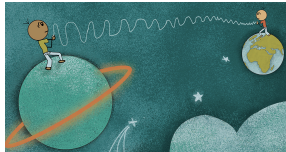
ENTANGLEMENT ENTROPY

In the $c \rightarrow +\infty$ limit, we have expansion

$$\log \mathcal{Z}_{\text{grav}}[\Lambda, c] \approx \vartheta + \left(\frac{c}{6} - \frac{19}{6} + \dots \right) \log \frac{1}{\Lambda} + \frac{c}{6} \log c + \dots$$

Logarithmic term \rightsquigarrow 2d CFT thermal/entanglement entropy

Holzhey,Larsen,Wilczek-...-Casini,Huerta-Cardy,Calabrese-...



Sub-leading terms \rightsquigarrow entanglement from gravity fluctuations

D.A.,Mühlmann

DEVELOPMENTS IN NEAR-DS₂ JT GRAVITY

- Sharp expressions for Hartle-Hawking WF as Schwarzian

Maldacena, Turiaci, Yang-Cotler, Jensen, Maloney-...

$$\Psi_{\text{HH}} = \mathcal{Z}_{\text{Schwarzian}}[L, \phi_b] = \left(\frac{\phi_b}{L}\right)^{3/2} e^{-i2\phi_b L + i\frac{4\pi^2 \phi_b}{L} + \frac{S_0}{2}} .$$

- 2d static patch with Dirichlet wall & near-Nariai geometry

Svesko, Verheijden, Verlinde, Visser-D.A., Harris-Jacobson, Visser-...

DEVELOPMENTS IN 3D GRAVITY WITH $\Lambda > 0$

- Novel solutions of black holes in dS_3 for modified Einstein's equation

$$\frac{ds^2}{\ell^2} = - \left(1 - r^2 - \frac{\mu}{r}\right) dt^2 + \left(1 - r^2 - \frac{\mu}{r}\right)^{-1} dr^2 + r^2 d\varphi^2$$

Empanan, Pedraza, Svesko, Tomasevic, Visser

- Novel Chern-Simons methods for quantum matter contributions to $\mathcal{Z}_{\text{matter}}$

$$\log \mathcal{Z}_{\text{scalar}} = i \int_{\mathcal{C}} \frac{d\alpha \cos \alpha/2}{\alpha \sin \alpha/2} \text{Tr}_j e^{\frac{\alpha}{2\pi} \oint A_L} \text{Tr}_j e^{\frac{\alpha}{2\pi} \oint A_R}$$

Castro, Coman, Fliss, Zukowski

- New ideas on dS_3/CFT_2 and WZW models

Hikida, Nishioka, Takayanagi, Taki

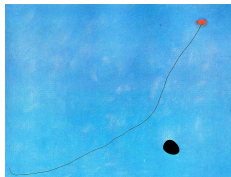
EMBEDDING IN A LARGER FRAMEWORK



WORLDLINE/WORLDTUBE PERSPECTIVE

de Sitter has no spatial boundary. Observations along timelike **worldline**?

...-D.A.,Hartnoll,Hofman-...-Witten-...



Naïvely problematic in gravity. Nonetheless, large N quantum mechanical models encode emergent space, e.g. BFSS, AdS₂/CFT₁

...-Polyakov...-Das,Jevicki-...-Banks,Fischler,Shenker,Susskind-...

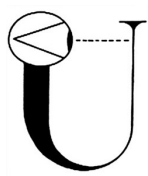
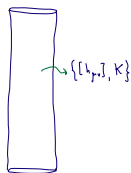
DE SITTER OBSERVER & TIMELIKE SURFACES

Decorate static patch with 'observer' to compute **von Neumann entropy** for resulting type II_1 algebra with maximally entropic state.

Chandrasekaran, Longo, Penington, Witten-..., building from: Leutheusser, Liu

Well-posedness for manifolds with timelike boundaries \rightsquigarrow fixing conformal class of induced metric + trace of extrinsic curvature.

Anderson-An, Anderson-Witten-Strominger, Bredberg-D.A., Anous, Bredberg, Ng-...-D.A., Galante, Mühlmann-...



Worldtubes as holographic surfaces in static patch?

York-...-Banks, Fischler-Susskind-...-D.A., Anous, Bredberg, Ng-...-Jacobson, Banihashemi-...-Blackler, Hartnoll

DYNAMICAL FEATURES OF DE SITTER HORIZON

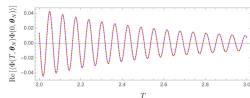
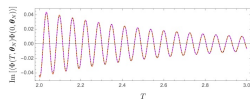
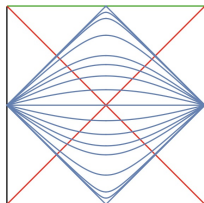
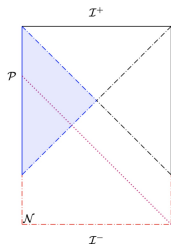
- Energy pulse \rightsquigarrow Horizon 'opens-up' more vertical Penrose diagram

Gao,Wald-...-D.A.,Hofman,Galante-...

- No (real) geodesics connecting spacelike separated points

...Galante,Chapman,Kramer-D.A.,Galante,Mühlmann-Chapman,Galante,Harris,Sheorey,Vegh-Aalsma,Faruk,van der Schaar,Visser,de Witte-...

- Oscillatory correlations for heavy particles & suppressed decay for light fields



TIMELIKE SURFACES IN dS_3

- Generalisations of $T\bar{T}$ -deformations \rightsquigarrow dS_3 physics with timelike boundaries

Coleman, Mazenc, Shyam, Silverstein, Soni-Shyam-...

- Opens up avenue to implement AdS_3/CFT_2 methods for dS_3 static patch

$$E_{\text{Brown-York}}[\ell/\ell_{Pl}, r_c] = E_{T\bar{T}}[c, \lambda]$$

building from: Smirnov, Zamolodchikov-Cavaglia, Negro, Szecsenyi, Tateo-...-McGough, Mezei, Verlinde-...

- Matching of logarithmic correction to Gibbons-Hawking entropy

EMBED DS_2 INTO AdS_2/CFT_1 FRAMEWORK

Simple model: Embed dS_2 in asymptotically AdS_2 world (deformed JT gravity)

D.A., Hofman-...-Ecker, Grumiller, McNees building from: Guth-Farhi-...-Freivogel, Hubeny, Maloney, Myers, Rangamani, Shenker-...-



Probe de Sitter horizon with **sharp observables** anchored at AdS_2 boundary.

TOWARD A MICROSCOPIC CONSTRUCTION

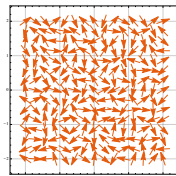
Progress in AdS_2/CT_1 duality \rightsquigarrow microscopic model

Sachdev, Ye-Kitaev-Rosenhaus, Polchinski-Maldacena, Stanford-...

dS as IR regime of SYK + relevant $\text{SL}(2, \mathbb{R})$ deformation

$$H = j_{i_1, \dots, i_q} \psi_{i_1} \dots \psi_{i_q} + \lambda j_{i_1, \dots, i_{q'}} \psi_{i_1} \dots \psi_{i_{q'}}$$

j are random variables & $i = 1, 2, \dots, N$



Analytically & numerically tractable **thermal renormalisation group flows**

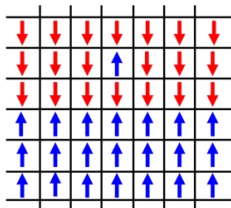
...-D.A., Galante, Sheorey-...

OUTLOOK



ISING MODEL FOR QUANTUM DE SITTER?

Guided by holographic paradigm & Euclidean methodology, **idealised models** & **concrete calculables** for quantum de Sitter space are beginning to surface



CONSTRAINTS ON MICROPHYSICAL TOY MODELS

- Universal features of $\mathcal{Z}_{\text{grav}}$

Reminiscent of [Sen's](#) logarithmic BPS black hole corrections. Perhaps 3d/2d gravity loop expansions for $\mathcal{Z}_{\text{grav}}$ can be reproduced by a microphysical model?

Coleman, Mazenc, Shyam, Silverstein, Soni-Mühlmann-Hertog, Bobev, Hong-...

- SYK inspired models

Some similar features between AdS_2 & dS_2 , suggesting [SYK](#) as a model. Basic challenge is to distinguish black hole/Rindler like features from dS features, e.g. radially shrinking horizon size...

D.A., Hofman-...-D.A., Galante, Sheorey-... also: -Susskind-Susskind, Lin-...

- Spectral constraints for $\mathcal{Z}_{\text{grav}}$

Analogy is [Klebanov-Giombi](#) sum of one-loop AdS higher spin tower reproducing dual $O(N)$ vector model. Negative edge modes produce novel source of cancellation. Engineer perturbative dS spectra exhibiting UV cancellations? SUSY inspired spectra?

D.A., BenettiGenolini, Mühlmann-D.A., RiosFukelman-...

QUANTUM INFORMATION IN QUANTUM COSMOLOGY?

- **Topology & wormholes in de Sitter inconclusive**

...-Chen,Gorbenko,Maldacena-...

$$\begin{aligned} \text{Tr } \rho_R^2 &= \text{Diagram 1} + \text{Diagram 2} \\ &= e^{-S_{\text{Hawking}}/\hbar} + e^{-S_{\text{Wormhole}}/\hbar} \end{aligned}$$

- **Absence of classical maximin surfaces**

...-Shaghoulian-Levine,Shaghoulian-...-Bousso,Penington-

- **Coupling to open/non-gravitational system may help clarify issues**

...-Hartman,Jiang,Shaghoulian-Aalsma,AguilarGutierrez,Sybesma-AguilarGutierrez,ChatwinDavies,Hertog,PinzaniFokeeva,Robinson-KamesKing,Verheijden,Verlinde-...

PRECISION COSMOLOGY

As we enter the era of precision **experimental** cosmology, we are also prompted into an era of precision **theoretical** cosmology

Quantitative data at **quantum level** bring us closer to clarifying quantities such as **de Sitter entropy**

Microphysical degrees of freedom of exponentially expanding spacetimes & dynamical properties are, likely, crucial to understand

No clear **consolidation/synthesis** of ideas yet, but common threads emerging

MUCHAS GRACIAS!

REVIEW OF REVIEWS

- Strominger, Spradlin, Volovich, “Les Houches Lectures on de Sitter Space” hep-th/0110007
- Witten, “Quantum gravity in de Sitter space” hep-th/0106109
- Bousso, “TASI Lectures on the Cosmological Constant” 0708.4231
- Anninos, “de Sitter Musings” 1205.3855
- Galante, “Modave Lectures on de Sitter space & Holography” 2304.xxxxx